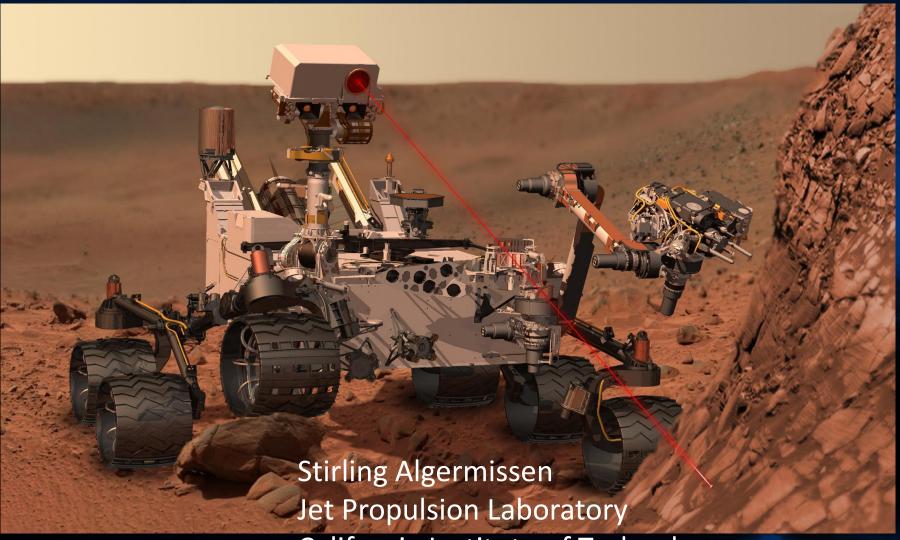


A Day in the Life of a Mars Rover Driver

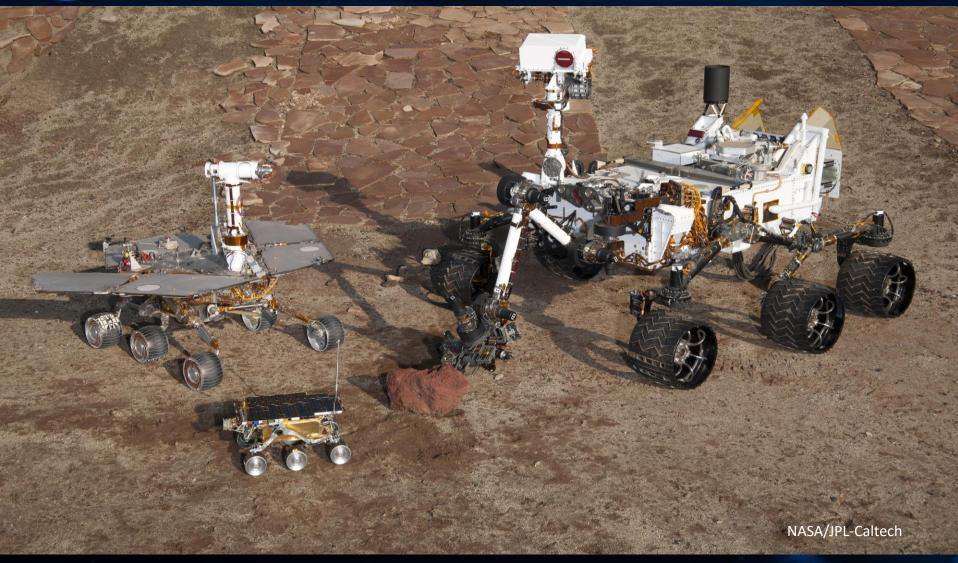


California Institute of Technology

Artist's Concept. NASA/JPL-Caltech



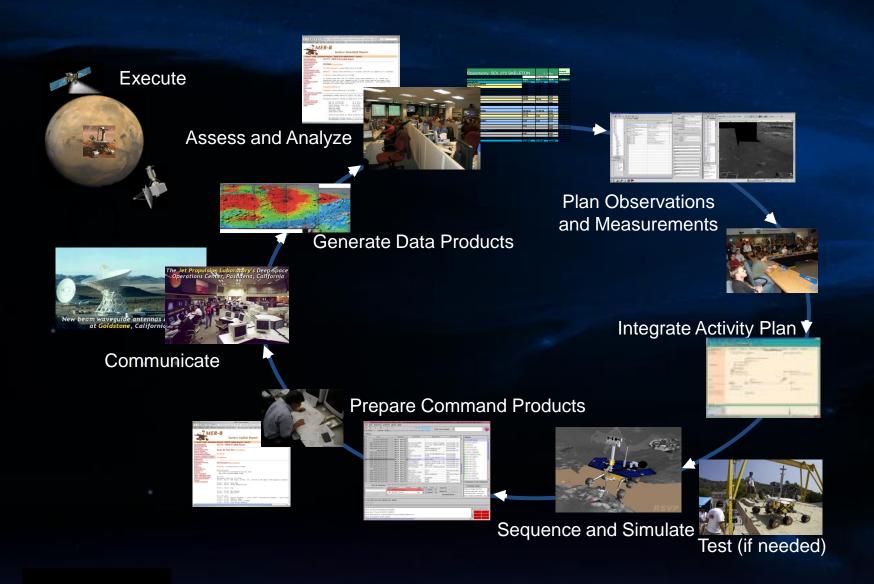
Mars Rover Family Portrait





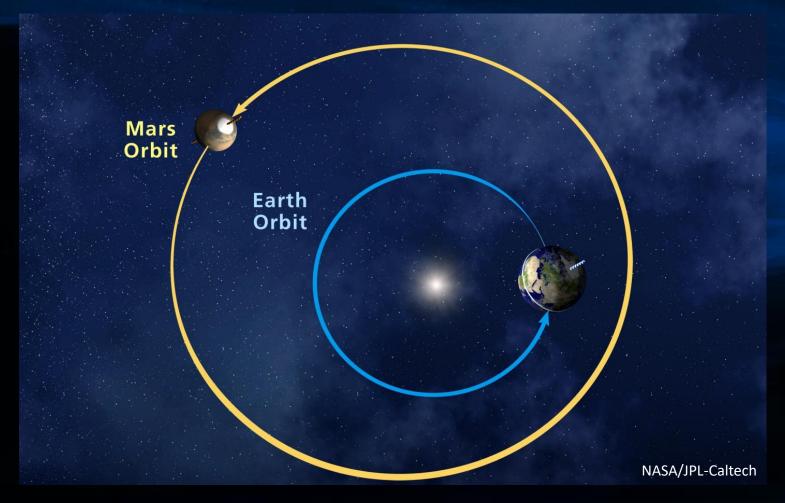


Mission Planning Cycle





Because of the distance between Earth and Mars, we can't drive a rover in real time.



It takes between 4 and 22 minutes each way for a signal to travel between the two planets.



Also, Logistics



The Deep Space Network is a shared resource for dozens of missions.

We often only get one uplink and few downlink windows each day



A previous day's images are fed into the Rover Simulation Visualization Program (RSVP) and 3D meshes are created.

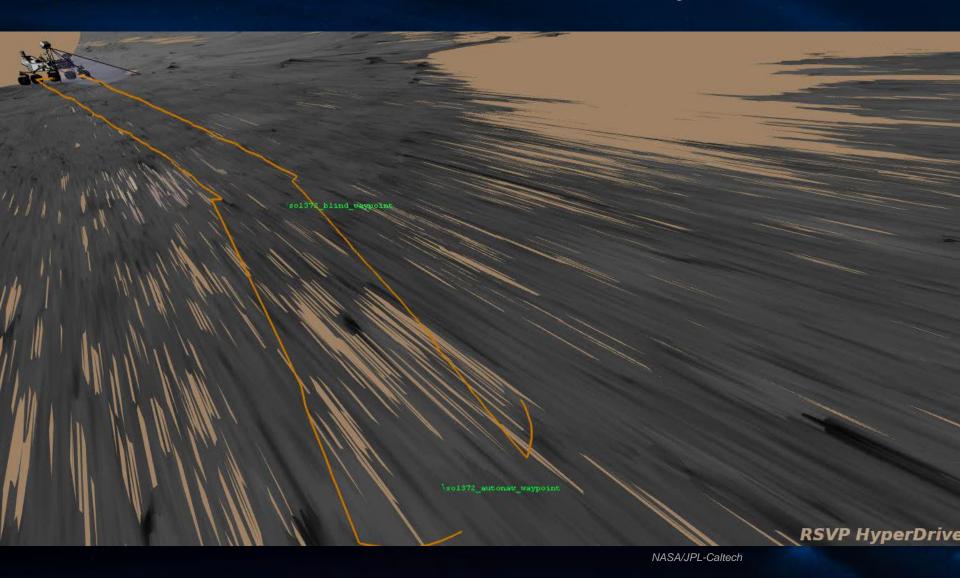


Rover drivers wear shuttered 3D goggles to view stereo imagery and 3D meshes

7

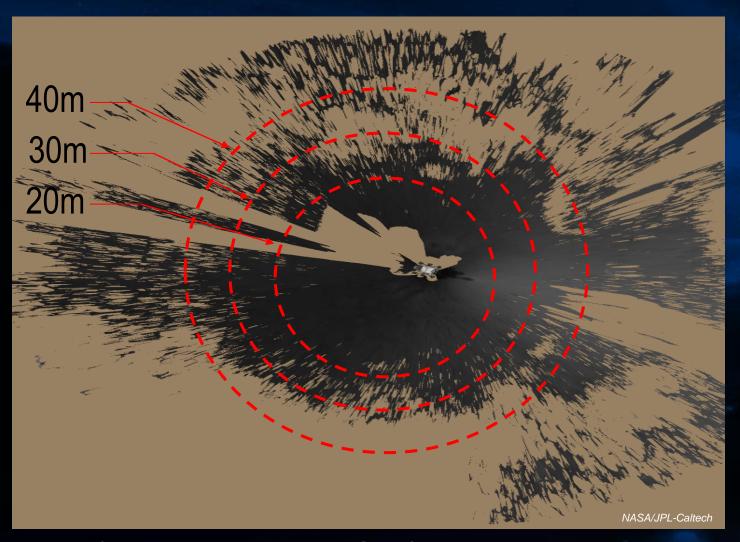


Drives are planned in 3D meshes out to the limit of visibility





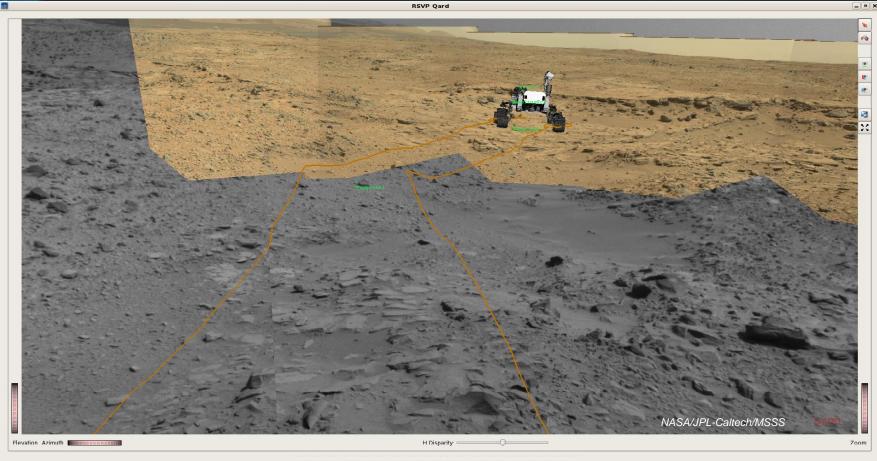
This image shows how much detail the Navcam cameras can typically see nearby.



3D data from Navcam stereo is often supplemented by color texture information in Mastcam images



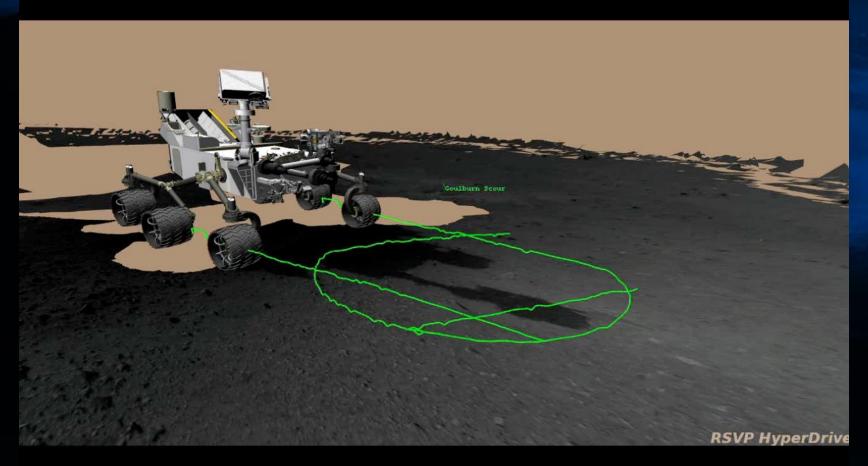
Multiple Drive Views



Drives are simulated on 3D meshes, and can be visualized in many ways. Here a mosaic of Navcam and Mastcam images provide context.



For "directed driving," drivers command the rover to move a certain distance over ground that they know is safe.



This is the fastest way to drive, because no predictive hazard processing is done, but distance is limited by what people can see. Curiosity will always stop the drive if a fault is detected!



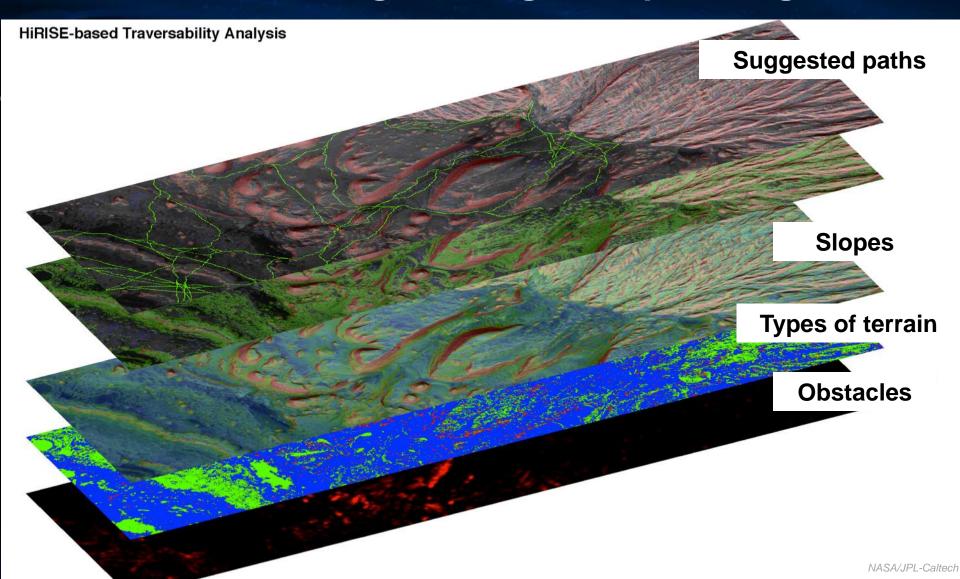
Curiosity carries out the activities and then sends data to the orbiters, whose larger antennas relay it to Earth.

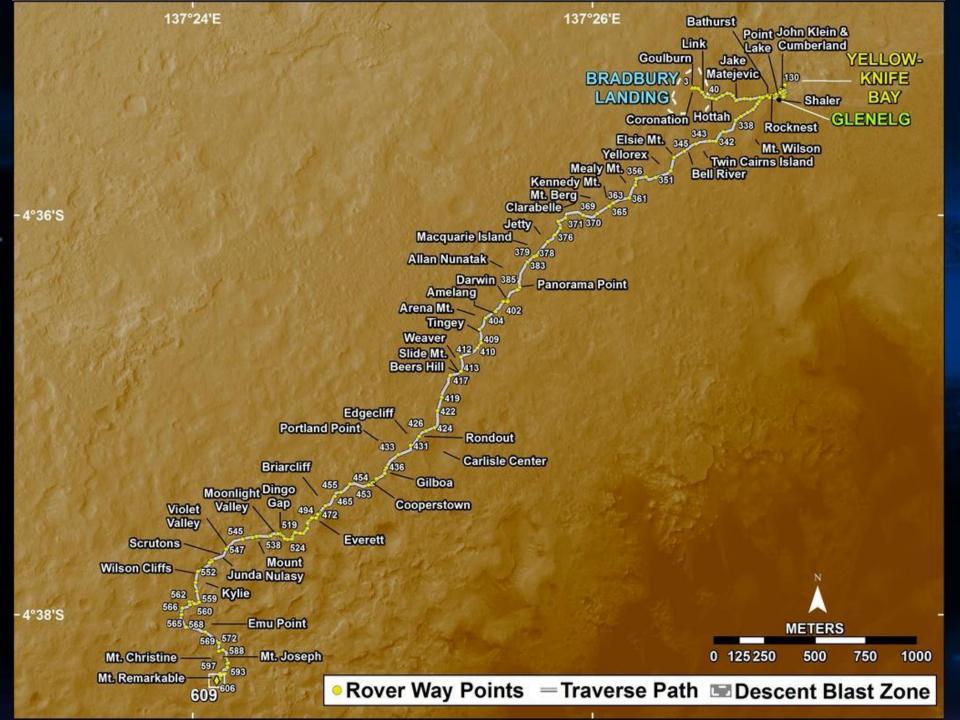


It takes less energy and a smaller antenna to send data 200 miles (322 km) up to an orbiter, rather than millions of miles to Earth, though direct contact is available.



Data from the Mars Reconnaissance Orbiter helps "see" several kilometers ahead, allowing for long term planning.





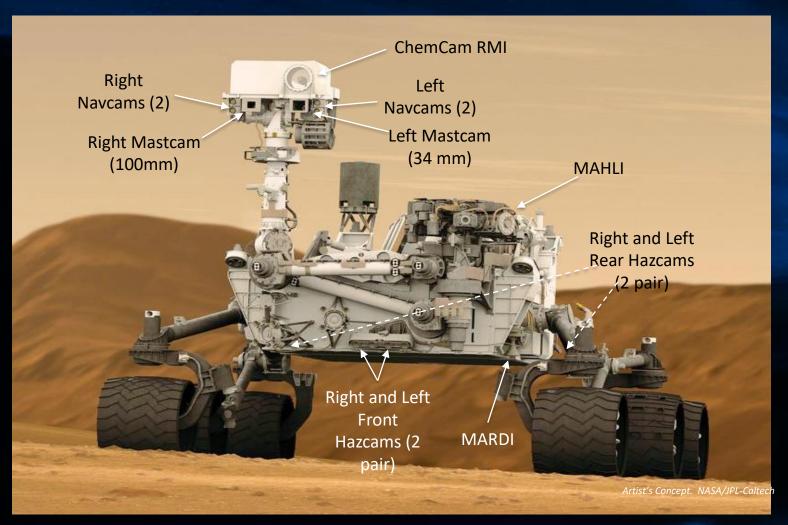


WDC? Robotics Tech!

Velocity-controlled Driving Autonomous Fault Response Visual Odometry Dense Stereo Vision Autonomous Terrain Assessment Local and Global Waypoint Planning Multi-sol Driving Visual Target Tracking Precision Arm Placement Percussive Drill **Cached Sample Manipulation Simulation Rover Sequencing and Visualization**

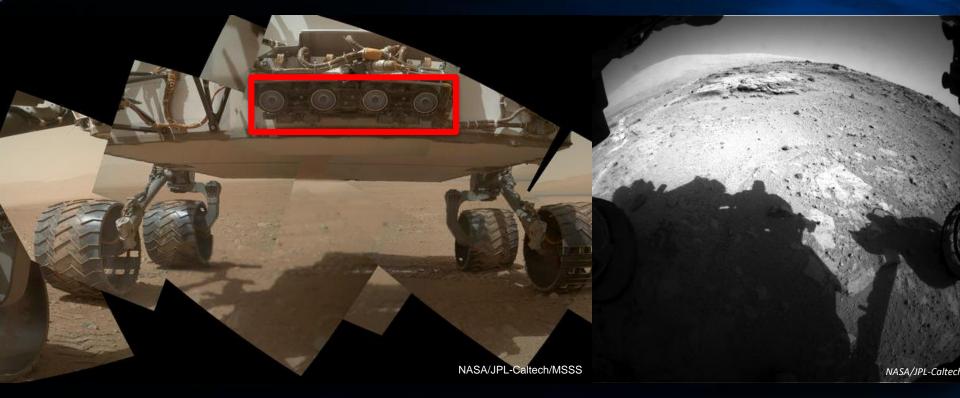


Curiosity has 17 cameras



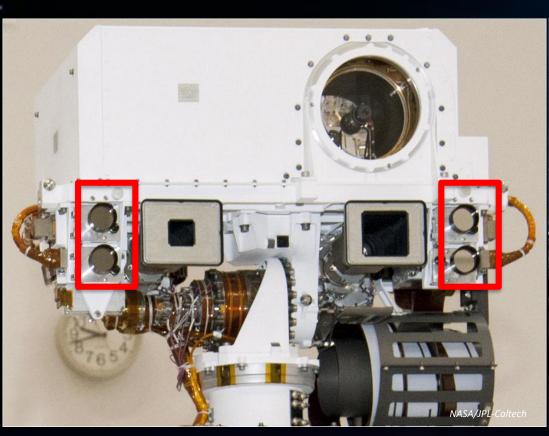
However, only the Hazcams and Navcams are tied into the auto-nav software.

The hazard avoidance cameras give a 120° wide angle view of the area near the rover. Front cameras have 16cm baseline, rear cameras have 10cm baseline.





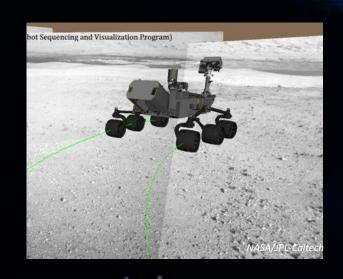
The 45° navigation cameras are almost 7 feet off the ground with 42cm baseline, providing good views over nearby obstacles or hills and into ditches.

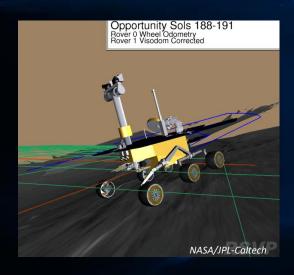


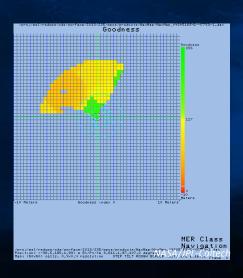




Human Rover Drivers Decide How Much Autonomy is Desired Based on Terrain and Available Resources







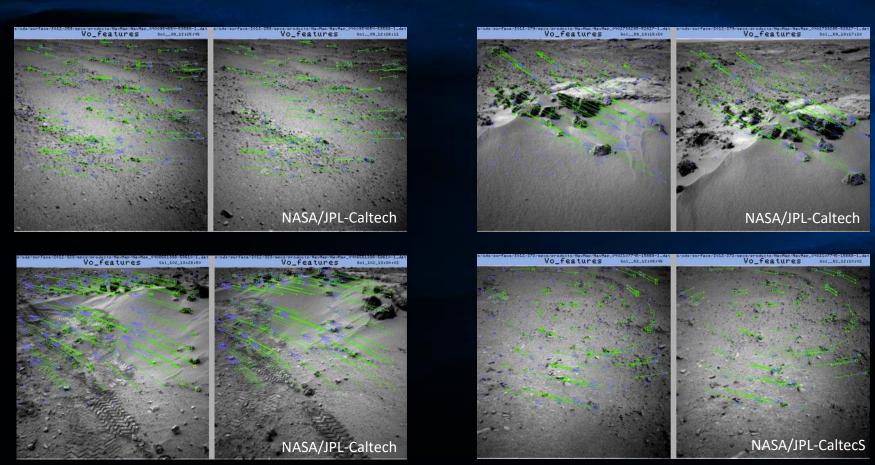
Directed driving

Visual odometry, or Slip Check + "Auto"

Auto-navigation; Geometric Hazard Detection and Avoidance



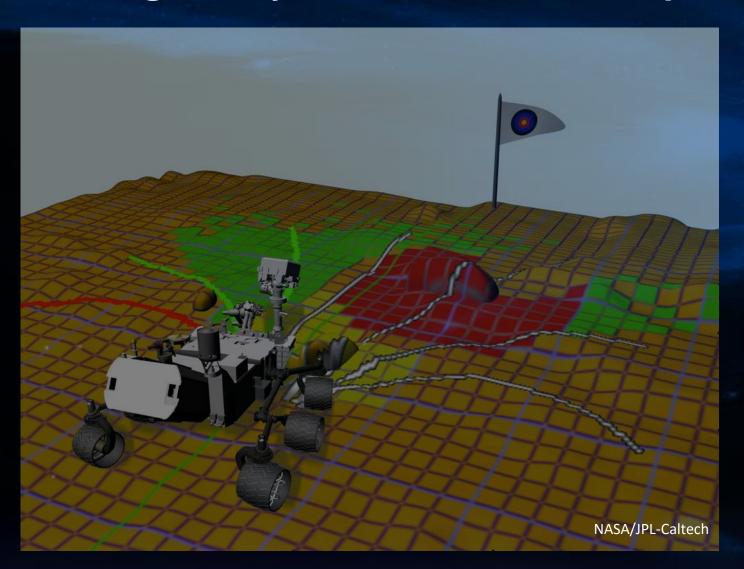
Using visual odometry, the rover constantly compares pairs of images of nearby terrain to calculate its position.



Unlike terrestrial robots, Curiosity drives as far as possible between VO images

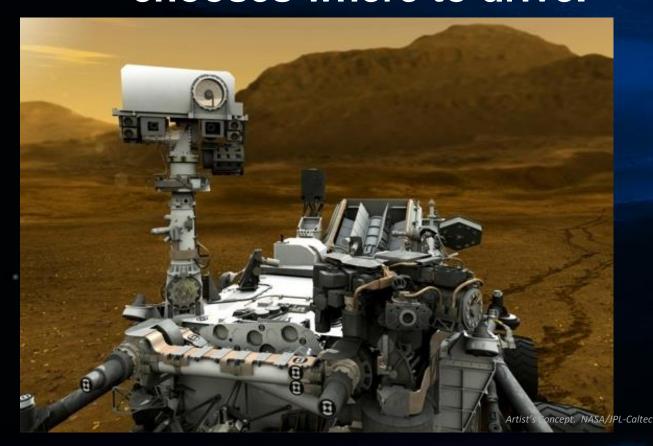


Human drivers and Curiosity depend on 3D image analysis to find the safest path.





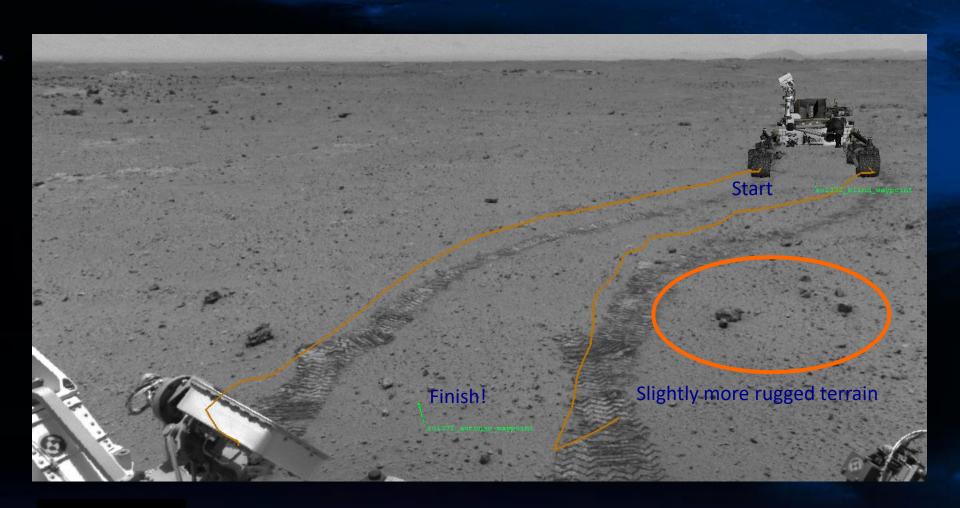
During nominal auto-nav, the rover stops every 0.5-1.5 meters, takes 4 sets of images, evaluates hazards, and then chooses where to drive.



Auto-nav extends directed drives into previously unseen terrain

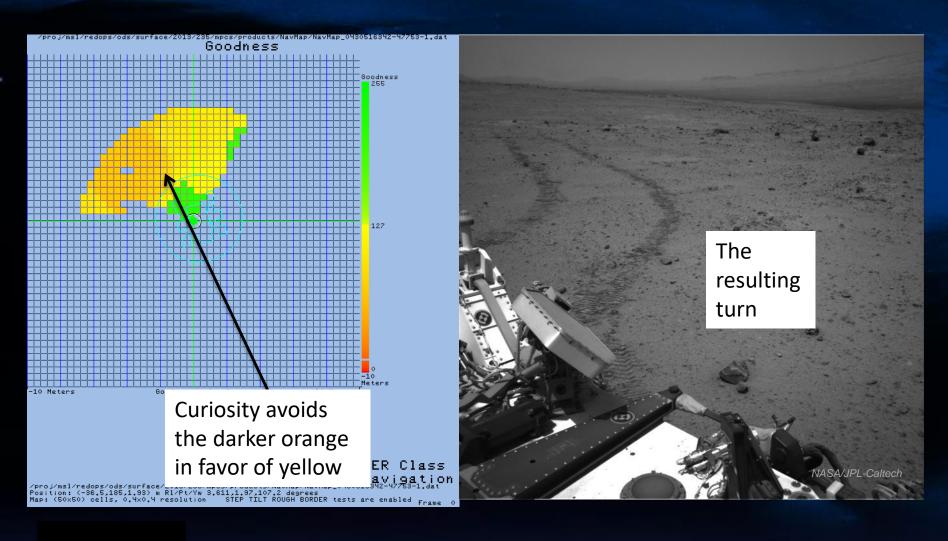


Wheel tracks after the first auto-nav drive on sol 372 show that Curiosity chose to drive around a little mound of loose rock.

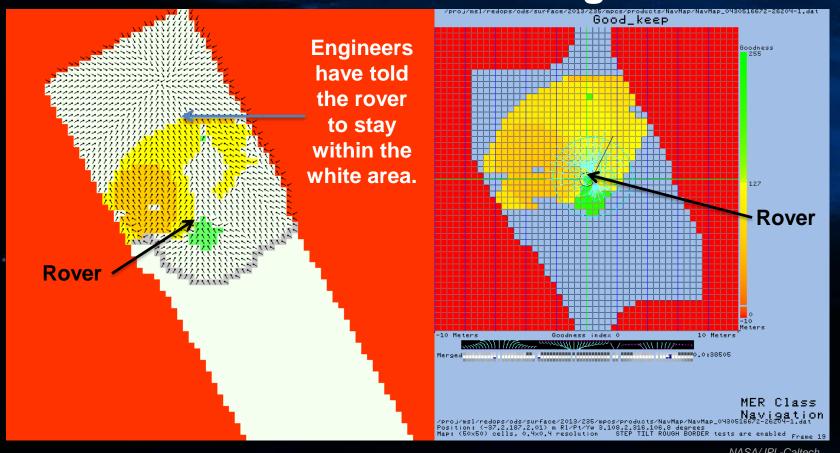




Curiosity's map and tracks show this decision to turn was based on her evaluation of the terrain.



The rover reduces a stereo point cloud into a configuration space, labeling unsafe areas red and safe areas green.

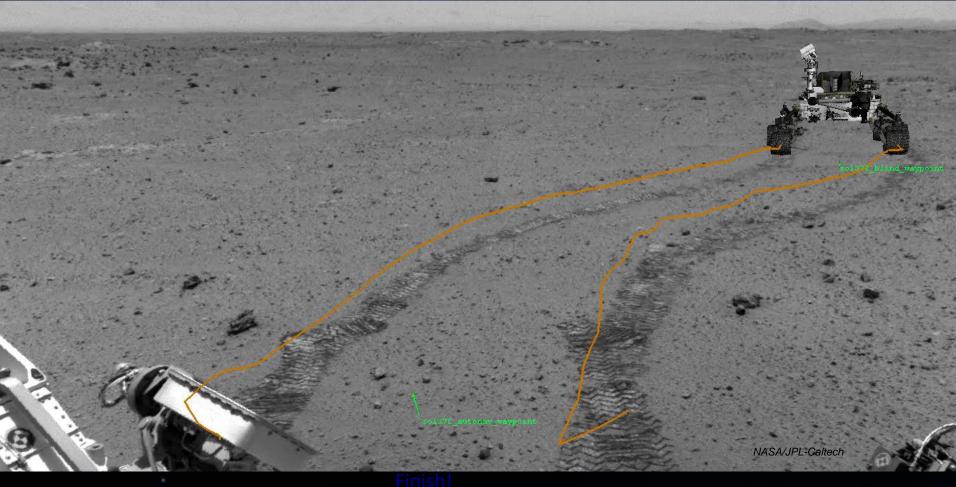


NASA/JPL-Caltech

Yellow means drive carefully, just like on Earth.

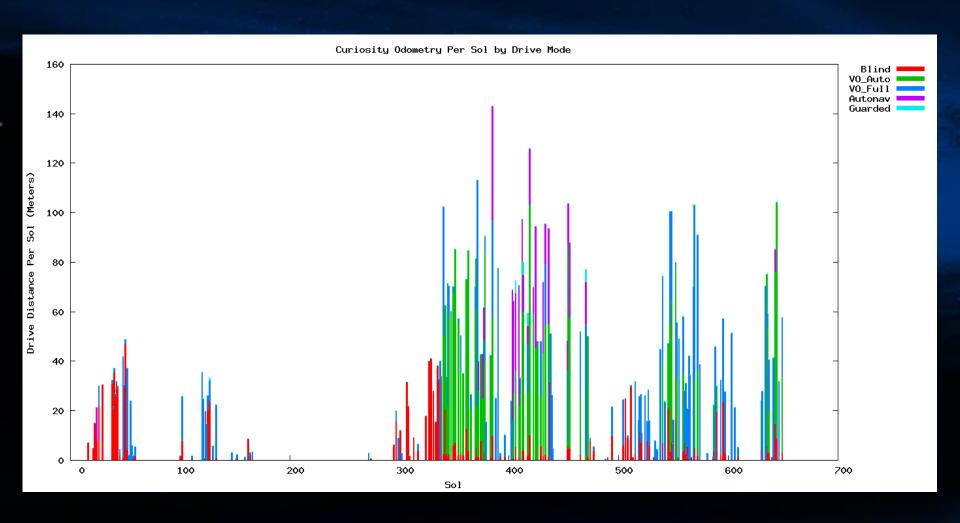


Animation of Curiosity's actual Sol 372 drive over a picture of her tracks



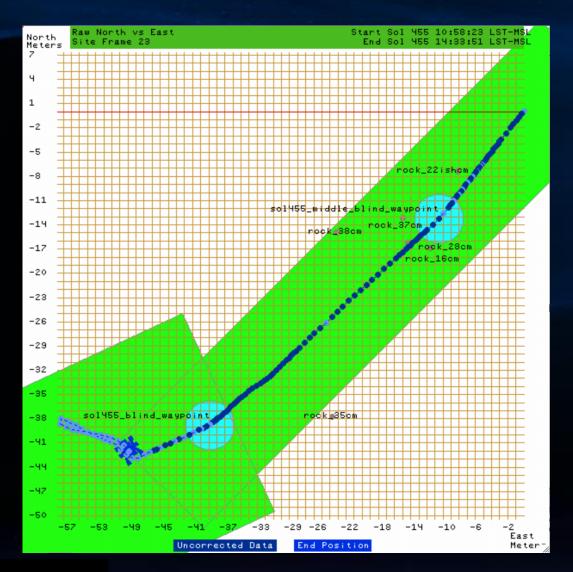


Curiosity Odometry Per Sol





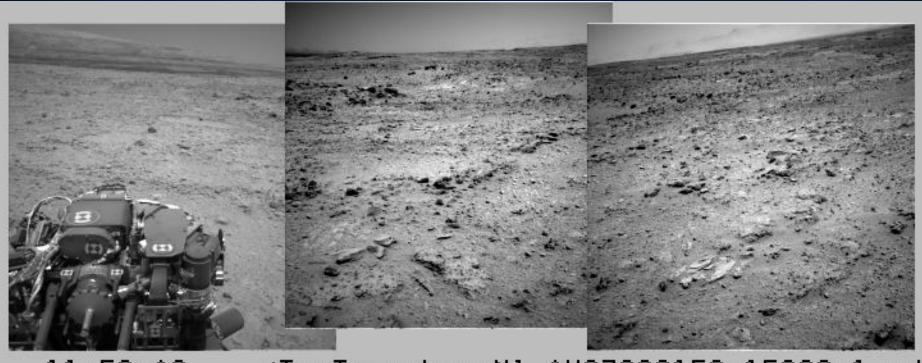
On sol 455, Curiosity encountered a small crater and began to drive around it



Small light blue dots represent the imaging steps



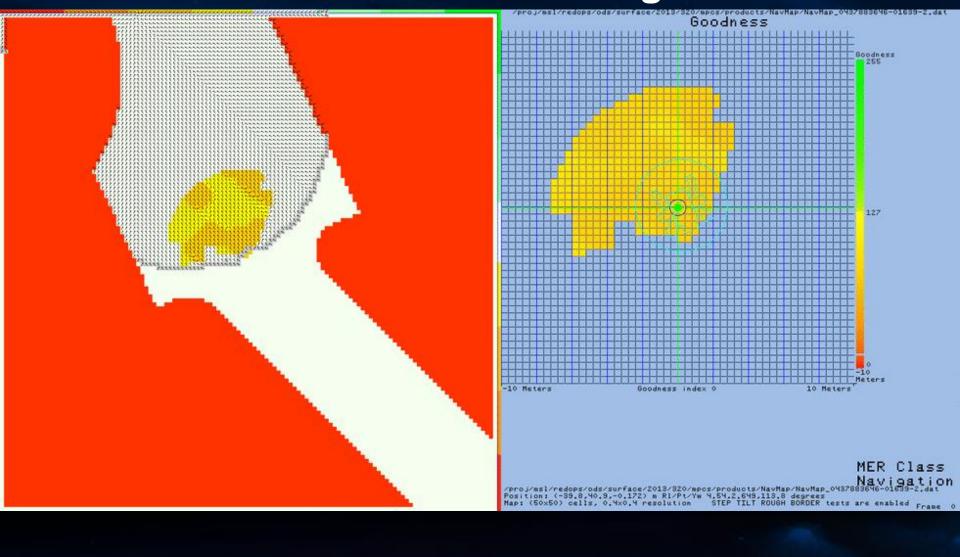
A Rover's-eye view of the Autonomous Portion of the sol 455 drive



11:59:02___./ImgImageLocoN1_0437883156-15288-1.pds

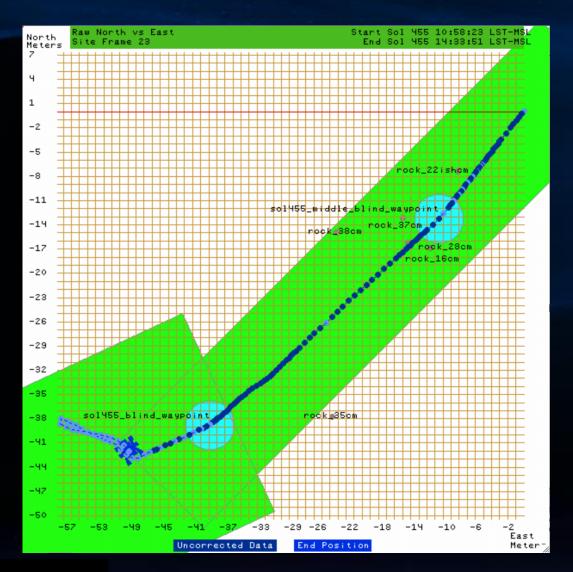


Then, boxed in by Keepin Zones, D* tried backtracking!





On sol 455, Curiosity encountered a small crater and began to drive around it



Small light blue dots represent the imaging steps



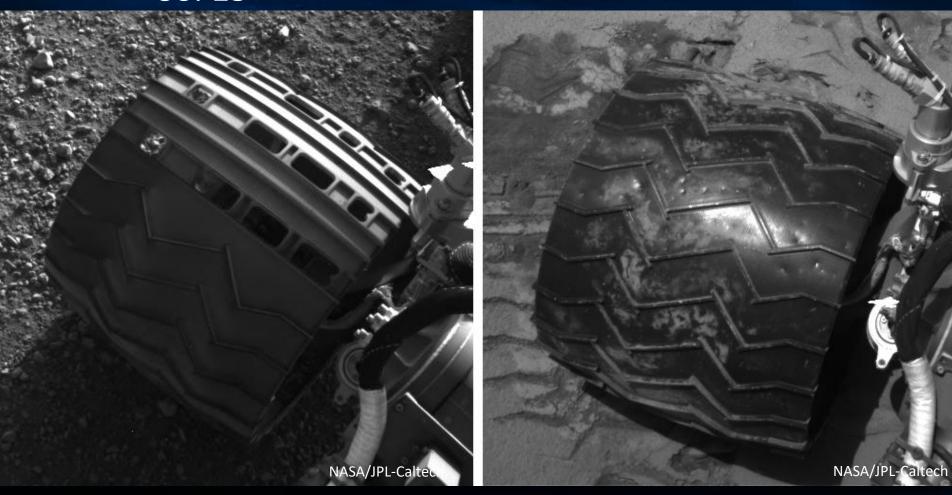
Sol 465: Wheel Wear

We started to notice unusual amount of wear in our wheels





Sol 15-59: Before and After Sol 15



Minor denting by small sharp rocks



Sols 60-309: Before and After

Sol 60 Sol 296



(No driving between sols 60-99, sols 167-271 and other smaller time windows)



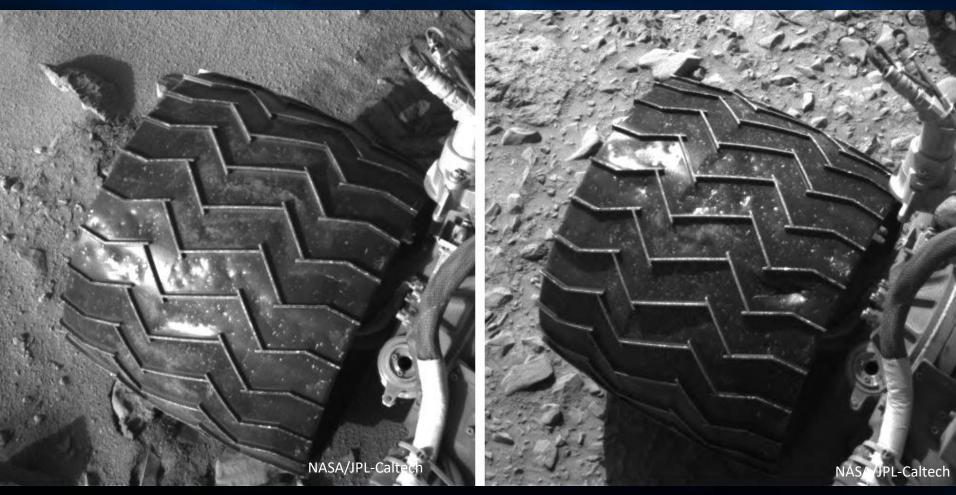
Sols 313-396: Before and After Sol 402 Sol 297



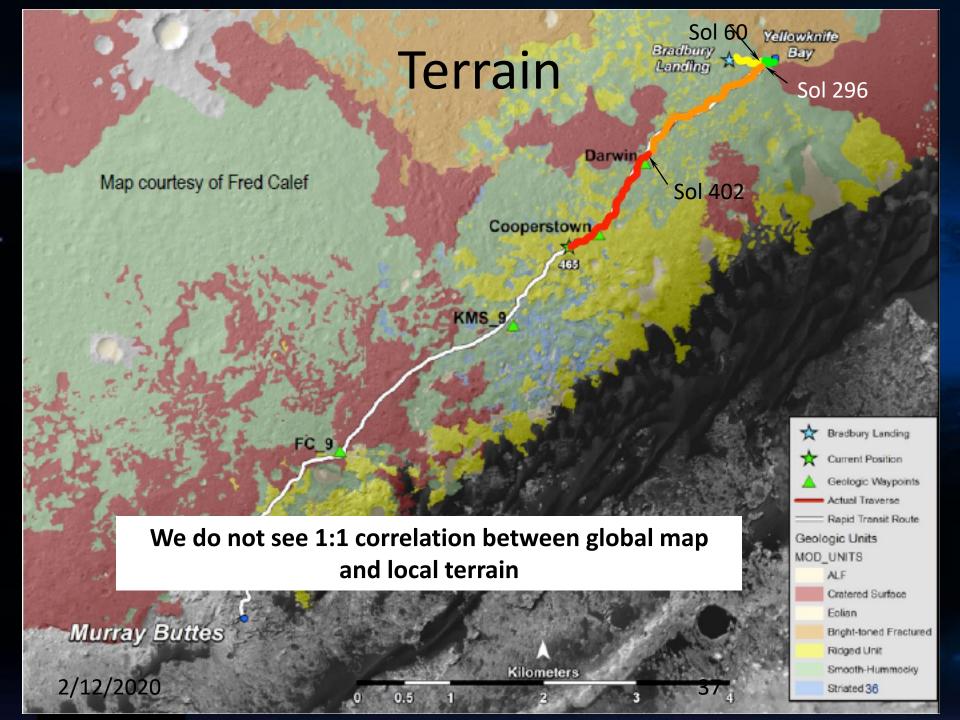
(No driving between sols 310-312 and sols 397-401; Wheel in partial image at sol 313 looks like that at sol 297)



Sols 402-477: Before and After Sol 402

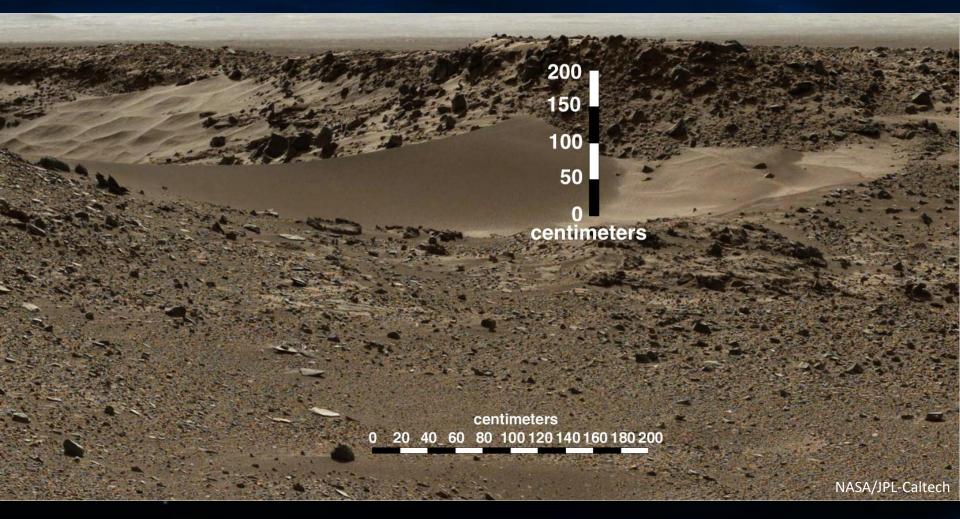


Severe denting by grouser-sized rocks



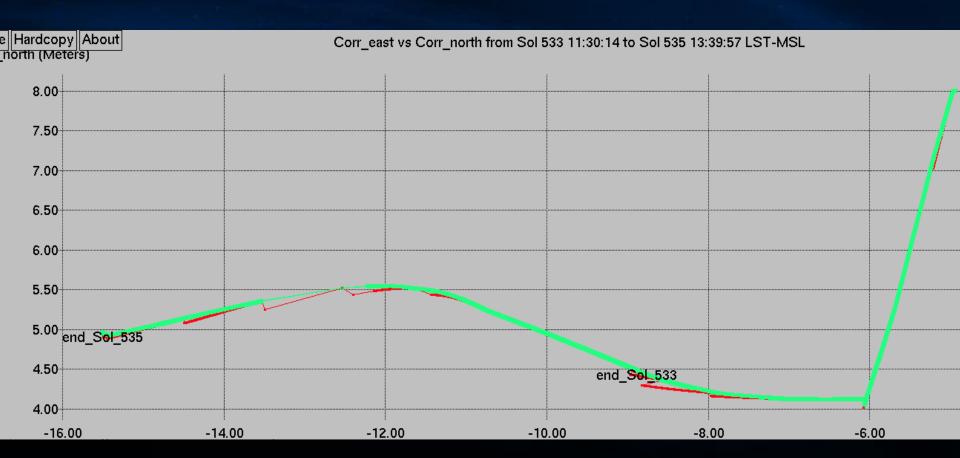


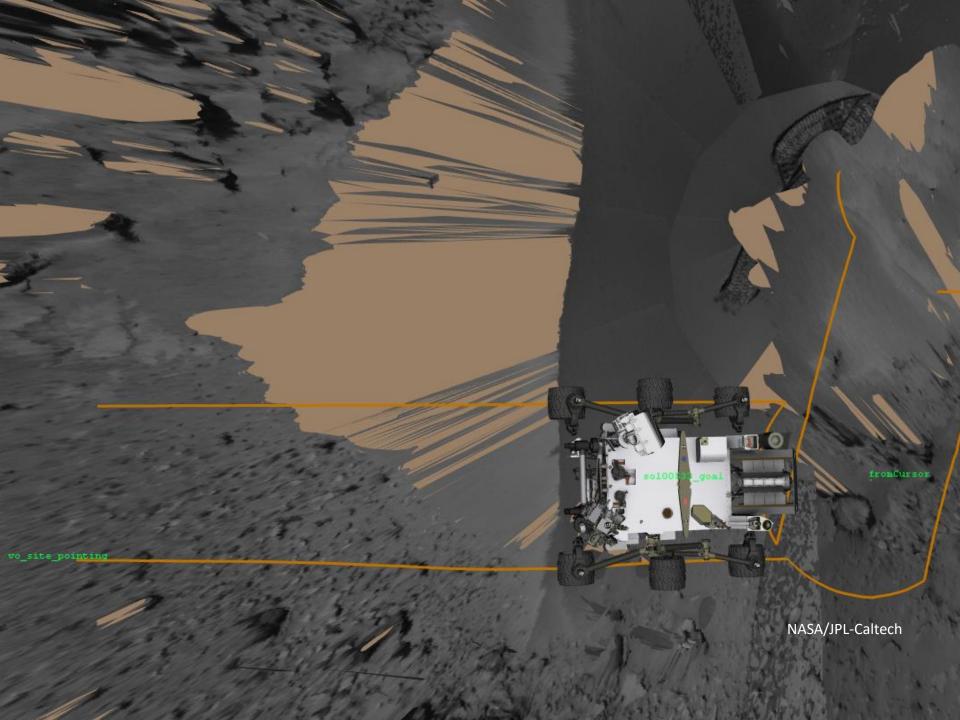
Sol 533-535: Dingo Gap





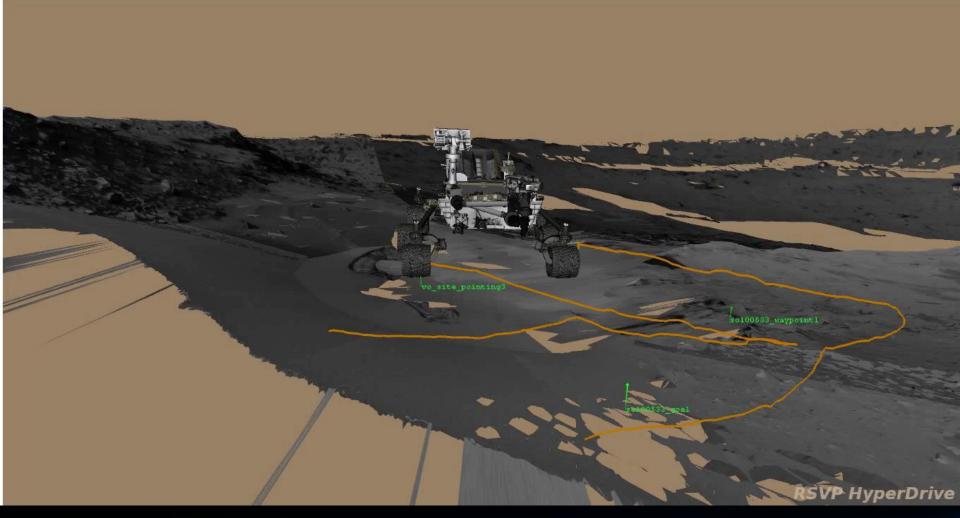
Sol 533-535: Dingo Gap







Sol 535: Climbing Over



NASA/JPL-Caltech

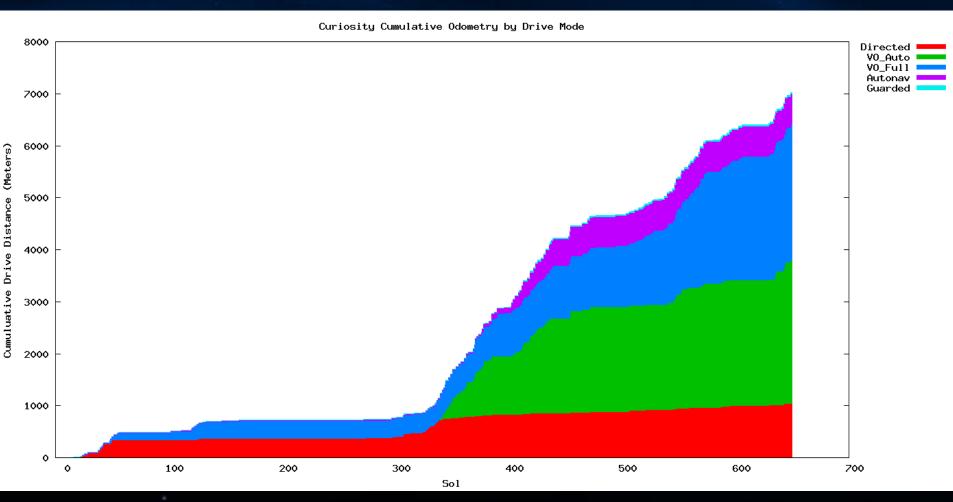


Sol 533-535: Dingo Gap



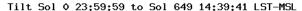


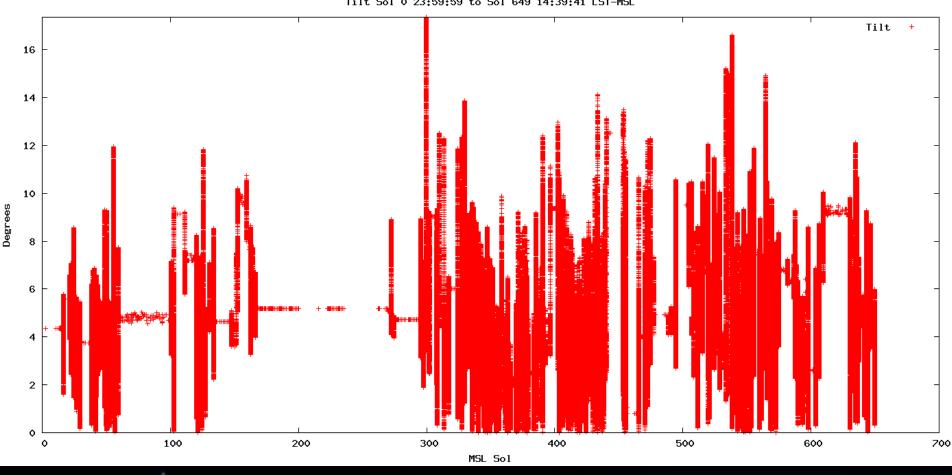
Curiosity Cumulative Odometry





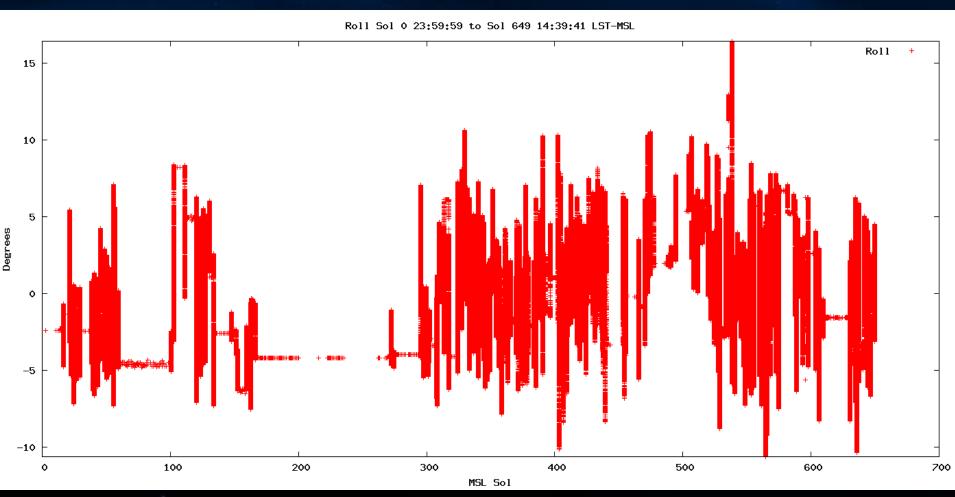
Curiosity Tilt Per Sol (through Sol 650)





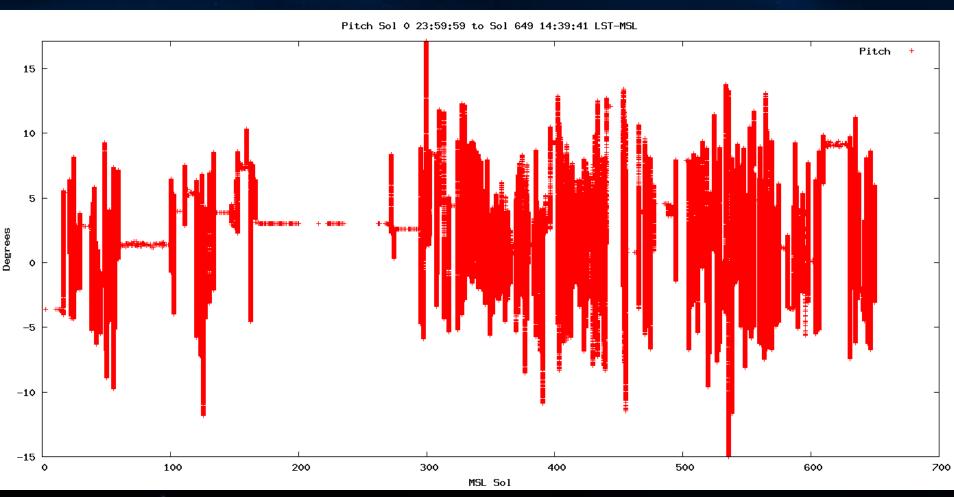


Curiosity Roll through Sol 650



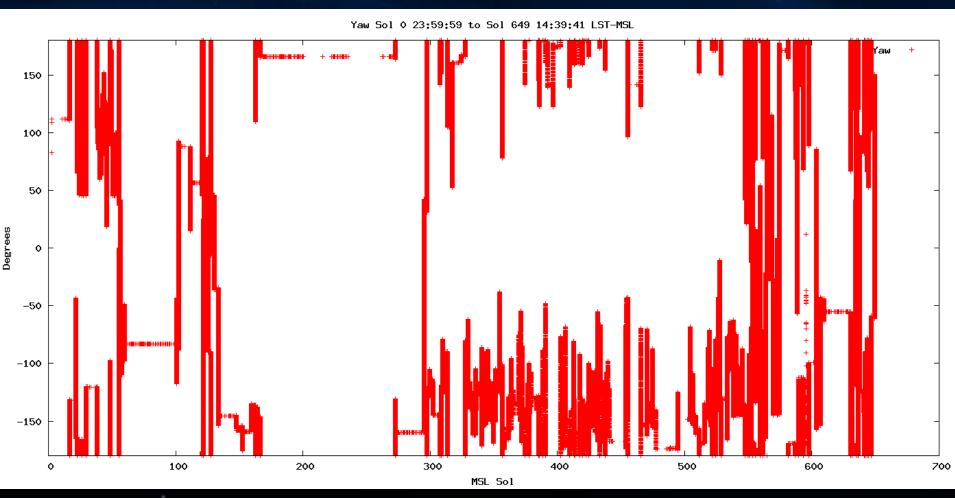


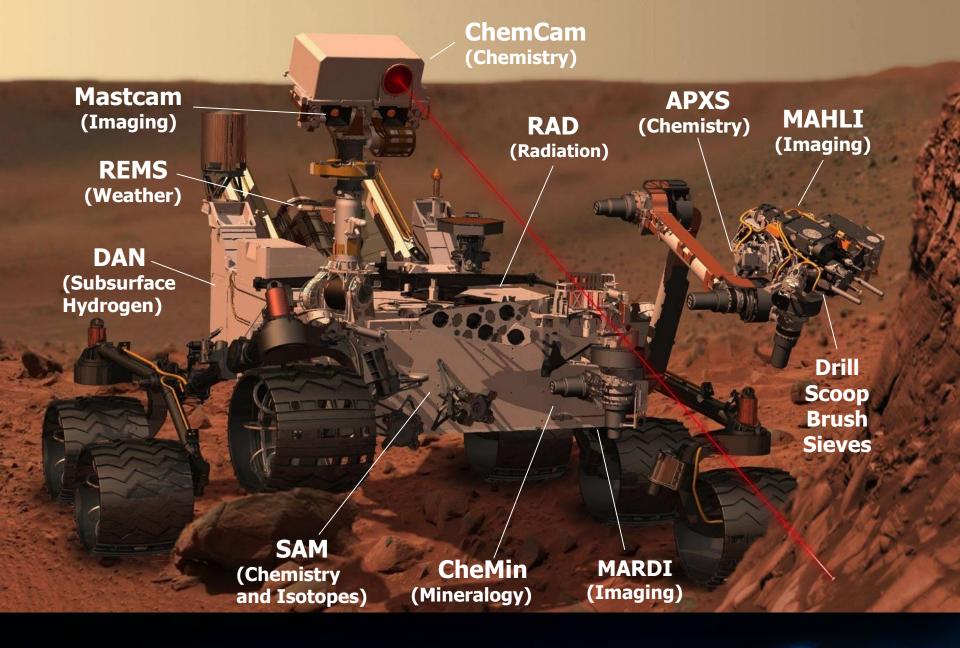
Curiosity Pitch through Sol 650



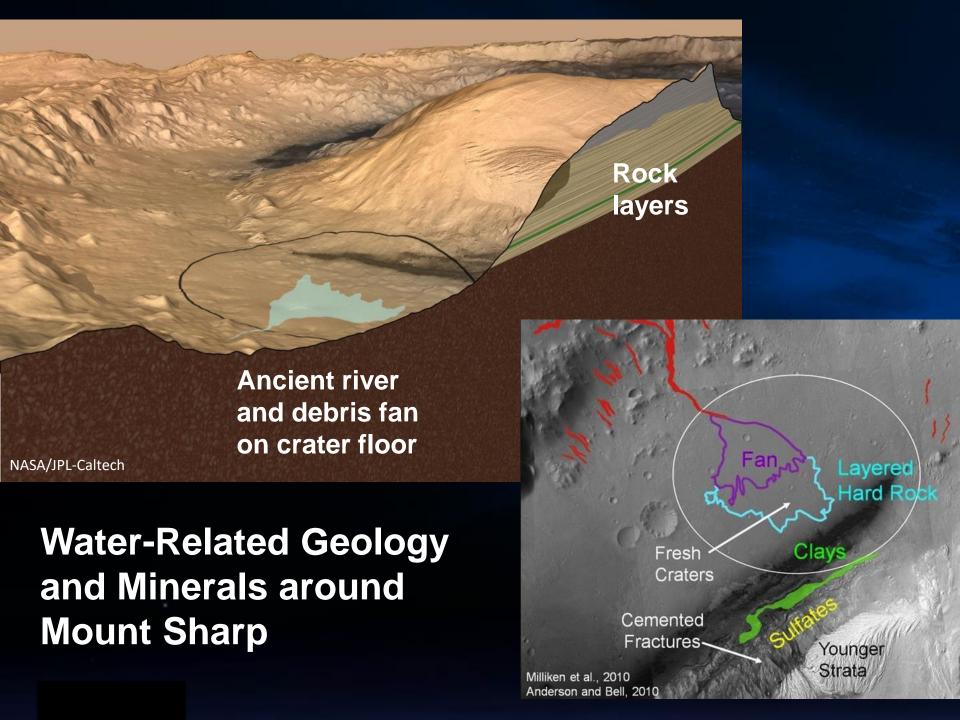


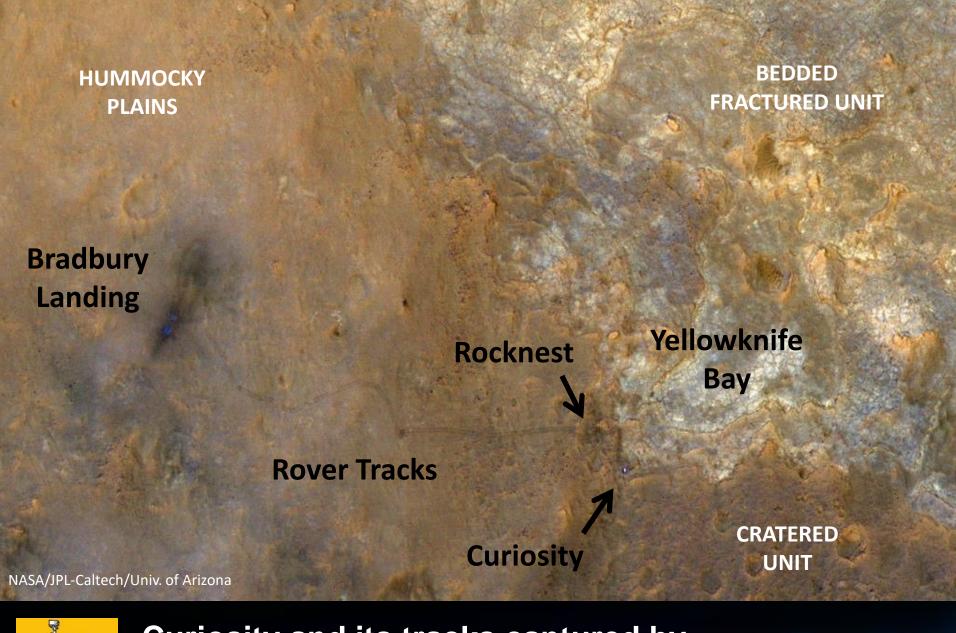
Curiosity Yaw through Sol 650





Curiosity's Science Payload Artist's Concept. NASA/JPL-Collector





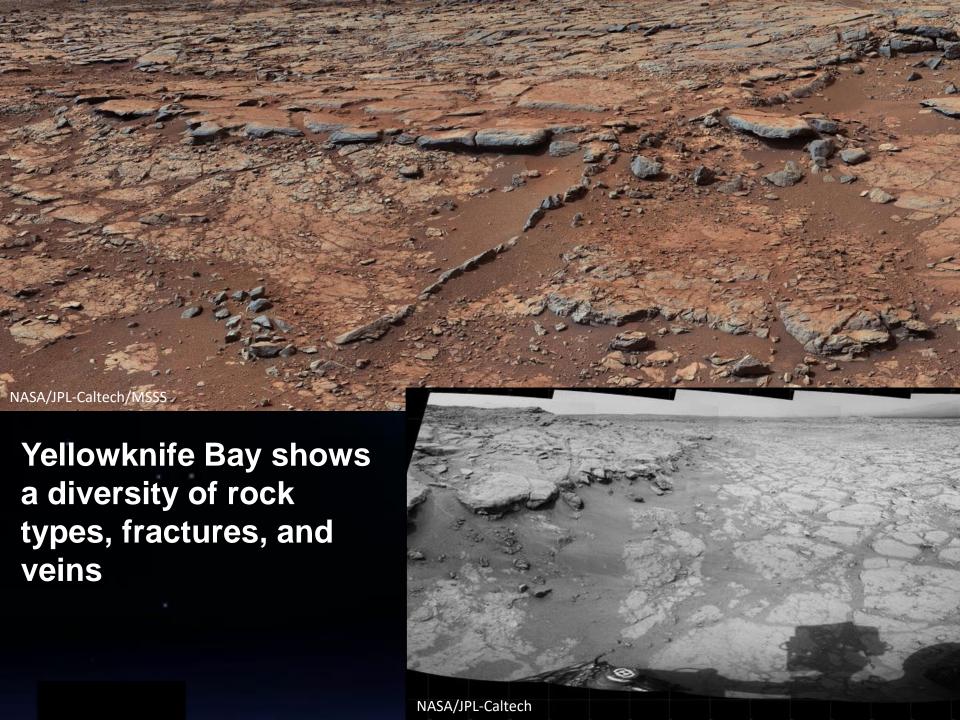


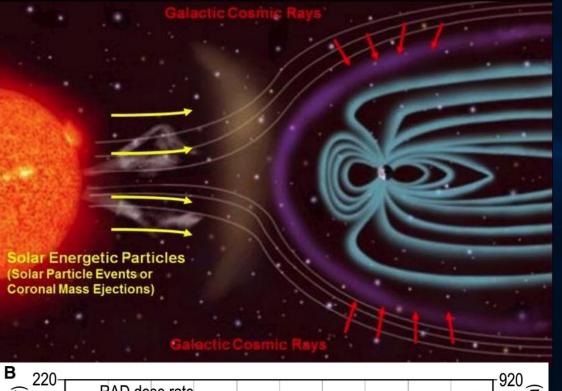
Curiosity and its tracks captured by HiRISE on the Mars Reconnaissance Orbiter





Rounded pebbles and sand in the conglomerate "Link" indicate water flowed ankle to hip deep





Jose Rate (micrcoGy/day) RAD dose rate 930 (G 215-**REMS Pressure** Pressure 210-205 760 720 680 (tuospheric) 200-195 190-640 185-22.000 23.000 24.000 25.000 21.000 26.000 Sols Since Landing

The RAD instrument measured the radiation flux from both galactic cosmic rays and solar energetic particles, in cruise and at Mars' surface

The surface dose rate is about half that measured in cruise

A crewed mission would receive ~1 Sievert of exposure in a trip to Mars with 500 sols on the surface

[Hassler et al., 2014]

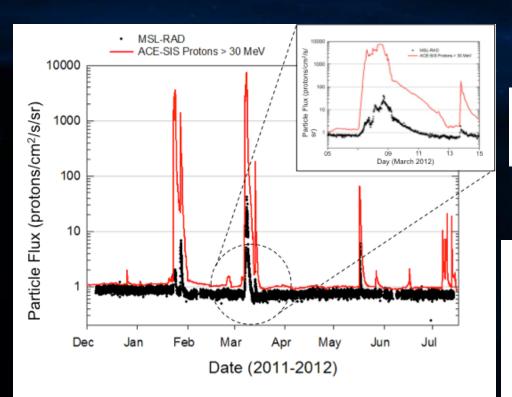
(Nep)
240
220
180
160
Sep Oct Nov Dec Jan Feb Mar Apr May Jun
Date (2012-2013)



Curiosity's Radiation Assessment Detector measures high-energy radiation



RAD & REMS



RAD observed galactic cosmic rays and five solar energetic particle events

RAD is now collecting the first measurements of the radiation environment on the surface of another planet

I get Mars weather reports from Twitter

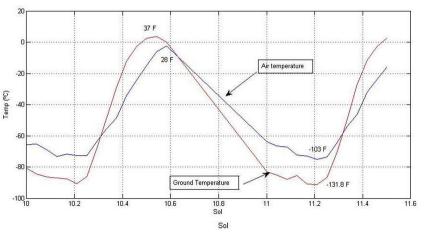




Mars Weather @MarsWxReport

Sol 76 (Oct 23, 2012): Sunny, high -1C/30F, low -72C/-97F, pressure higher at 7.91 hPa, wind E at 7.2kmh/4.5mph, daylight 6am-5pm Expand

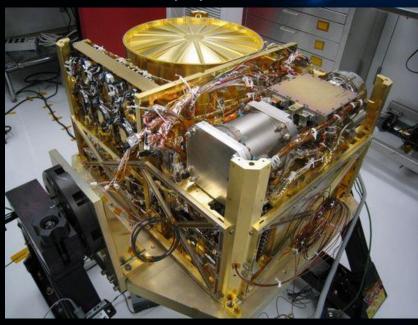
GROUND AND AIR TEMPERATURE SENSOR





SAM & CheMin

SAM instrument which takes up more than half the science payload on the rover

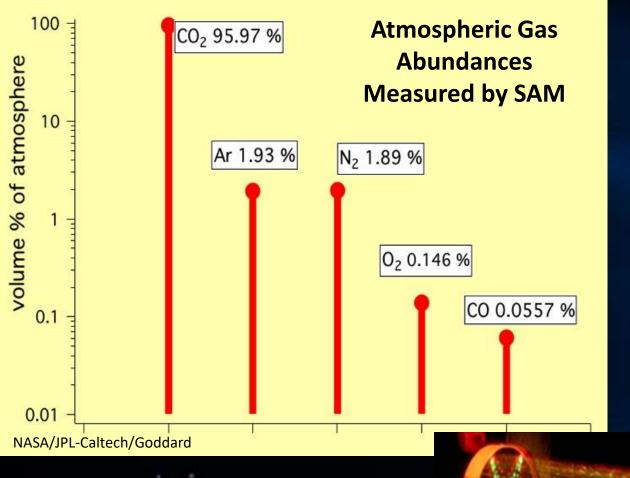


Sample Analysis at Mars (SAM) is the rover's Easy Bake Oven.

It heats soil and rock samples samples until they vaporize and then analyzes the resulting gases CheMin Inlet



CheMin uses X-rays to determine mineral content and crystal structure of surface samples



SAM found that argon, rather than nitrogen is the second most abundant gas

SAM also found that Mars' atmosphere is enriched in the heavy versions of isotopes, indicating massive atmospheric loss to space

$$\delta^{13}$$
C = 46 ± 4 per mil

$$\delta D = 4950 \pm 1080 \text{ per mil}$$

40
Ar/ 36 Ar = 1900 ± 300

Methane has not been definitively detected

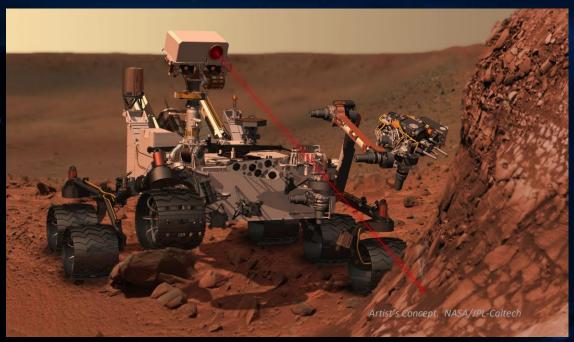
Upper limit = 1.3 ppb

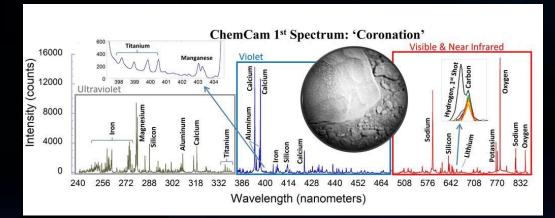


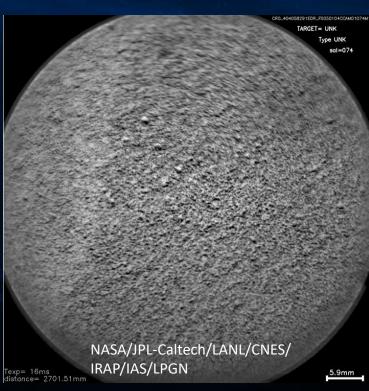
The SAM Tunable Laser Spectrometer and Mass Spectrometer measure atmospheric composition



A Laser (ChemCam)











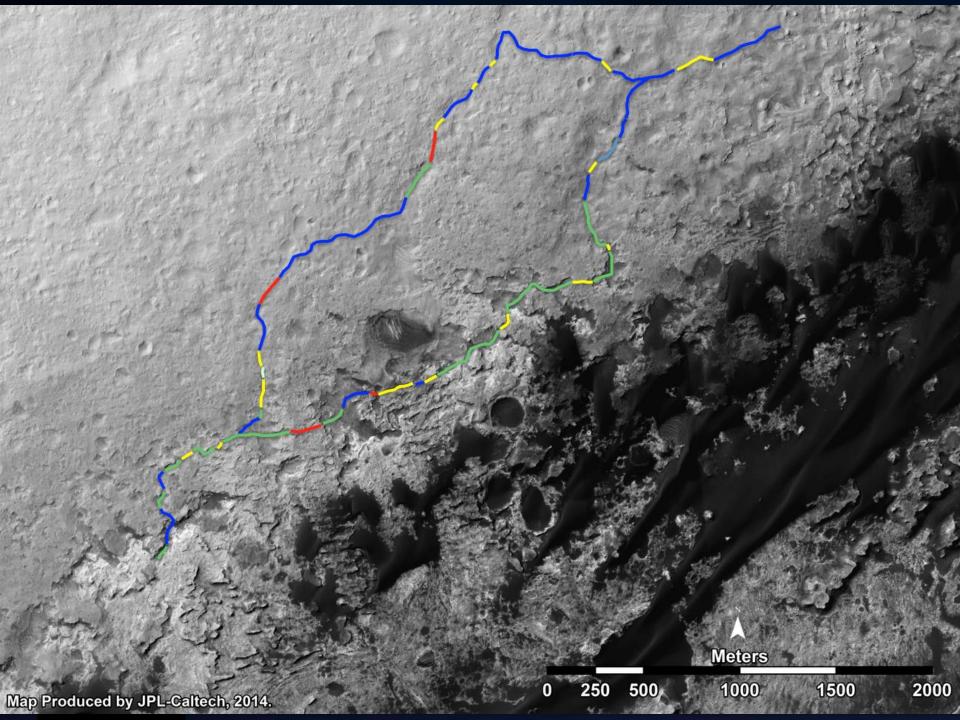
NASA/JPL-Caltech/MSSS



John Klein dime-sized drill hole with light-toned veins and ChemCam profile

An Ancient Habitable Environment at Yellowknife Bay

- The regional geology and fine-grained rock suggest that the John Klein site was at the end of an ancient river system or within an intermittently wet lake bed
- The mineralogy indicates sustained interaction with liquid water that was not too acidic or alkaline, and low salinity.
 Further, conditions were not strongly oxidizing.
- Key chemical ingredients for life are present, such as carbon, hydrogen, oxygen, phosphorus, and sulfur
- The presence of minerals in various states of oxidation would provide a source of energy for primitive organisms



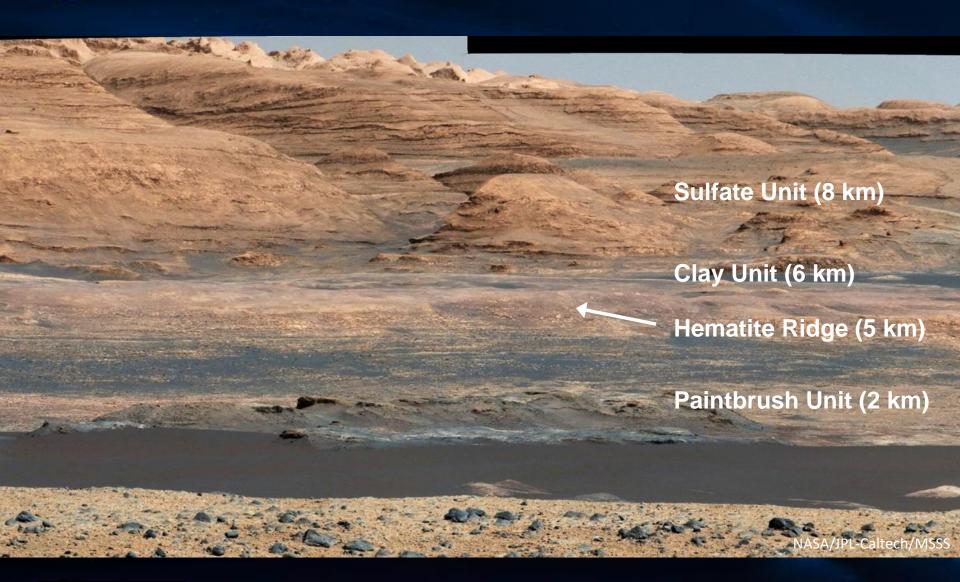


The Road Ahead

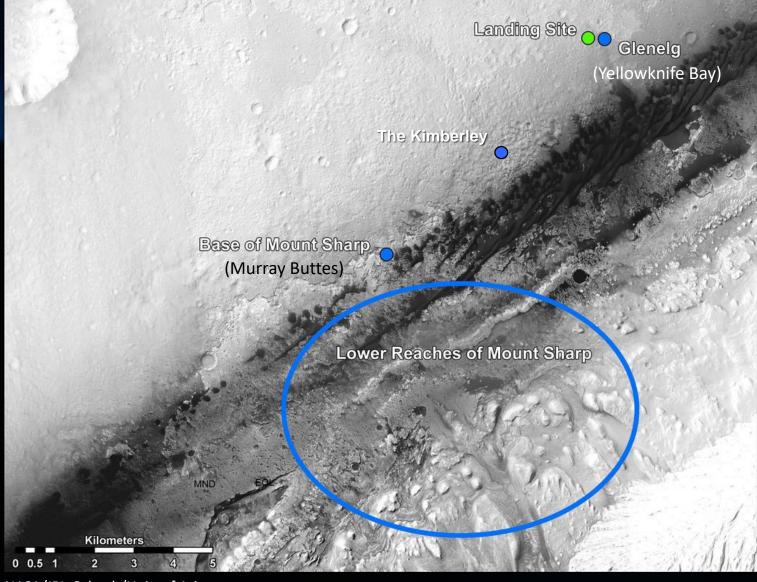




Targets for Exploration







NASA/JPL-Caltech/Univ. of Arizona

Curiosity's ultimate goal is to explore the lower reaches of the 5-km high Mt. Sharp

